

## Bulk sound travel times and implications for mantle composition and outer core heterogeneity

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**Abstract.** We present a new methodology for constraining mantle composition directly from observed seismic travel times. By measuring travel times of *PcP*, *PcS*, *ScS*, and *PKiKP* phases generated by the Chinese nuclear explosion of 21 May 1992 and recorded at epicentral distances of  $\sim 10^\circ$ , we directly determine bulk sound travel times ( $T^\Phi$ ) for the mantle and outer core. This eliminates the need to combine independent  $V_P$  and  $V_S$  profiles, characterizing different paths and frequencies, to obtain a reference  $V_\Phi$  profile. Our observed *PKiKP* - *PcP* differential travel time indicates that the outer core model of AK135 is slow by 2.1 s (that of IASP91 by 2.3 s) while that of PREM is slow by only 0.7 s and suggests, through comparison with other data, relative lateral homogeneity of the outer core. Furthermore, we find that  $T_0^\Phi$  for the crust and mantle in this region exceeds the AK135 prediction by 3.5 s (IASP91 by 3.9 s) +1.0 s. Since the lower mantle contribution to this delay depends upon how much arises in the upper mantle, we present two end-member models for such a slower mantle  $V_\Phi$ , in which the delay resides entirely above or below 660 km depth. The latter yields best-fitting lower mantle compositions which are less enriched in Si than those fit to  $T_0^\Phi$  for AK135 (or IASP91) for the former, and neither supports Fe-enrichment of the lower mantle. Applicability of such localized observations to the larger mantle is suggested by tomographic results indicating that lateral heterogeneity in  $V_\Phi$  is very small relative to that in  $V_S$ . Our results illustrate the potential utility of  $T_0^\Phi$  in constraining mantle composition.

### Introduction

Constraints upon the composition of Earth's interior are frequently derived by comparison of the radial profiles of density and seismic velocity obtained from inversion of seismic data to those obtained from various mineralogical models [Birch, 1952; Jackson, 1983; Knittle *et al.*, 1986; Jeanloz and Knittle, 1989; Bina and Silver, 1990; Stixrude *et al.*, 1992; Hemley *et al.*, 1992; Zhao and Anderson, 1994]. Direct comparison with seismological *P*-wave velocity ( $V_P$ ) and *S*-wave velocity ( $V_S$ ) profiles is difficult, since determination of  $V_P$  and  $V_S$  for mineral assemblages requires knowledge of the poorly constrained behavior of mineral shear moduli

at high temperatures and pressures. Therefore comparisons are frequently made using the bulk sound velocity ( $V_\Phi$ ), given by

$$V_\Phi^2 \equiv V_P^2 - \frac{4}{3}V_S^2, \quad (1)$$

in order to avoid the additional uncertainties associated with extrapolation of shear moduli [Bina and Silver, 1990]. However, construction of a seismological  $V_\Phi$  profile is complicated by the fact that the  $V_P$  and  $V_S$  profiles employed in equation (1) generally have been independently determined for different regions [Bina and Wood, 1987]. Thus, the component  $V_P$  and  $V_S$  profiles may be characteristic of different temperature structures, compositional regimes, and frequency ranges, leading to systematic errors in the resulting  $V_\Phi$  profile. Hence attempts to constrain mantle composition, for example, can be as sensitive to uncertainties in reference seismic models as to uncertainties in mineral thermoelastic equations of state [Bina and Silver, 1990].

Here we present a new methodology for directly determining bulk sound travel times ( $T^\Phi$ ) for the mantle and outer core in a single region, thus circumventing the need to combine independent  $V_P$  and  $V_S$  profiles characterizing different paths and frequencies. We use these  $T^\Phi$  measurements to test different models of outer core velocity and to evaluate the sensitivity of inversions for lower mantle chemistry to uncertainties in both seismic data analysis and mineral equations of state.

### Method

We measure the travel times of core-reflected phases recorded at small source-receiver epicentral distances ( $\Delta \sim 10^\circ$ ). In such experiments, core-reflected *P* and *S* waves are recorded on the same instruments, and they traverse nearly identical paths which are essentially independent of Earth model. For a given phase recorded at a distance  $\Delta$ , we measure the difference ( $\delta T_\Delta$ ) between the observed travel time and that predicted ( $T_\Delta$ ) by a reference model such as AK135 [Kennett *et al.*, 1995] (or IASP91 [Kennett and Engdahl, 1991]). We correct  $\delta T_\Delta$  to its value at  $\Delta = 0^\circ$  ( $\delta T_0$ ) by assuming

$$\frac{\delta T_0}{T_0} \approx \frac{\delta T_\Delta}{T_\Delta} \quad (2)$$

for small  $\Delta$ .

Given any two of the phases *PcP*, *PcS*, *ScS*, and *ScP*, we can determine two-way  $T_0^\Phi$  through the crust

and mantle by considering small variations in velocity at a given depth  $z$ :

$$(V_{\Phi} + \delta V_{\Phi})^2 = (V_P + \delta V_P)^2 - \frac{4}{3}(V_S + \delta V_S)^2. \quad (3)$$

Upon vertical integration over depth, this yields the perturbation in vertical bulk sound travel time ( $\delta T_0^{\Phi}$ ):

$$\delta T_0^{\Phi} = -2 \int_0^{CMB} \left[ \left(1 + \frac{4}{3}R\right)^{\frac{3}{2}} \frac{\delta V_P}{V_P^2} - \frac{4}{3}R^{\frac{3}{2}} \frac{\delta V_S}{V_S^2} \right] dz, \quad (4)$$

where  $R \equiv V_S^2/V_{\Phi}^2$ . If we assume that  $R$  is independent of depth through the mantle, this becomes:

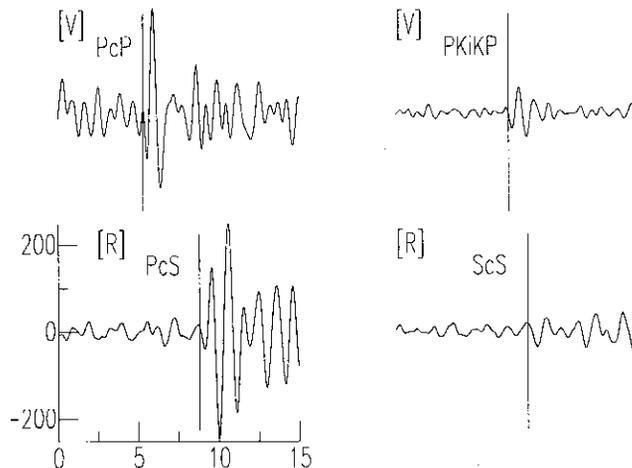
$$\delta T_0^{\Phi} \approx \left(1 + \frac{4}{3}R\right)^{\frac{3}{2}} \delta T_0^P - \frac{4}{3}R^{\frac{3}{2}} \delta T_0^S, \quad (5)$$

where  $\delta T_0^P$  and  $\delta T_0^S$  are the perturbations in vertical  $PcP$  and  $ScS$  travel times, respectively. The actual  $\pm 5\%$  variations of  $R$  about its mean for AK135 (0.4924) induce error in  $\delta T_0^{\Phi}$  of  $\pm 0.05$  s.

Given three or more of these phases, we can perform a consistency check by deriving several pairwise estimates of  $\delta T_0^{\Phi}$ , thus providing a simple measure of uncertainty.

## Data and Results

We use seismic waves generated by the Chinese nuclear explosion of 21 May 1992, recorded by a PASSCAL instrument deployment on the Tibetan Plateau [Owens *et al.*, 1993] and subjected to a 4-pole Butterworth bandpass filter of 0.75-2.00 Hz (Figure 1). Core-reflected phases arrive late relative to AK135 (and IASP91) predictions. Delays at SANG ( $\Delta = 10.80^\circ$ ,  $\phi_{az} = 166.62^\circ$ ) are 1.6, 1.8, and 1.1 s for  $PcP$ ,  $PcS$ , and  $ScS$ , respectively, and delays at USHU ( $\Delta = 10.77^\circ$ ,  $\phi_{az} = 140.04^\circ$ ) are 1.0 and 0.2 s for  $PcS$  and  $ScS$ , respectively, including corrections for ellipticity and for  $S$ -wave splitting due to known receiver-side anisotropy [McNamara *et al.*, 1994] at SANG ( $\Delta t = 0.80$  s,  $\phi_{az} =$



**Figure 1.** The core phases  $PcP$  (vertical),  $PKiKP$  (vertical),  $PcS$  (radial), and  $ScS$  (radial) recorded at SANG ( $\Delta = 10.80^\circ$ ,  $\phi_{az} = 166.62^\circ$ ) and bandpass filtered to 0.75-2.00 Hz. For each window, 15 seconds of data are shown.

$51^\circ$ ) and USHU ( $\Delta t = 0.72$  s,  $\phi_{az} = 119^\circ$ ). No  $ScP$  phases are observed at these stations.

Pairwise estimates of  $\delta T_0^{\Phi}$  relative to AK135 at SANG yield 2.9 s from  $PcP$  &  $ScS$ , 4.6 s from  $PcS$  &  $ScS$ , and 2.5 s from  $PcP$  &  $PcS$ . A pairwise estimate at USHU yields 3.9 s from  $PcS$  &  $ScS$ . Thus, observation of these phases yields  $T_0^{\Phi}$  for the crust and mantle which exceeds the AK135 prediction by 3.5 s (IASP91 by 3.9 s)  $\pm 1.0$  s, so that mantle  $V_{\Phi}$  in this region must be slower than AK135 (similarly for IASP91 and for PREM [Dziewonski and Anderson, 1981]). Reassuringly, our delay falls within the broad uncertainties of previous results from long-period multiple- $ScS$  phases [Sipkin and Revenaugh, 1994] which indicate that  $T_0^S$  for our receiver region is slow by  $1.3 \pm 1.4$  s while that for our source region is fast by  $0.8 \pm 4.3$  s.

Additionally,  $PKiKP$  is observed at SANG, constraining  $V_P (= V_{\Phi})$  for the outer core in this region. The observed differential travel time between  $PKiKP$  and  $PcP$  is 476.8 s, indicating that the AK135 outer core model is slow by 2.1 s (IASP91 by 2.3 s). By contrast, the PREM outer core is slow by only 0.7 s (presumably due in part to differing core model radii).

## Discussion and Conclusions

Our  $PKiKP - PcP$  differential time observation indicates that the PREM core model provides a better fit to actual core travel times beneath this region than does the AK135 (or IASP91) core model. Furthermore, the relevant differential times used to derive PREM (477.5 s at  $10.9^\circ$  and 477.2 s at  $11.73^\circ$ ) were recorded from NTS nuclear explosions in the western United States [Engdahl *et al.*, 1974], thus limiting to 0.7 s the vertically averaged difference in outer core velocities beneath these two widely separated regions and arguing for a laterally homogeneous outer core.

Our  $PcP$ ,  $PcS$ , and  $ScS$  observations can be used to constrain allowable lower mantle compositions. The structure of the crust and upper mantle (0-660 km depth) and that of the very top and bottom of the lower mantle (661-764 km, 2726-2889 km) are constrained by features other than  $T_0^{\Phi}$ . We therefore begin by fixing densities ( $\rho$ ) and velocities ( $V_{\Phi}$ ) for these regions at their AK135 values. By allowing the composition of the bulk of the lower mantle (765-2725 km) to vary in terms of Mg/Fe content and Si content, given by the molar ratios  $X_{Mg} \equiv Mg/(Mg + Fe)$  and  $X_{Si} \equiv Si/(Mg + Fe)$ , respectively, we solve for the lower mantle compositions whose calculated  $\rho(z)$  and  $V_{\Phi}(z)$  profiles along adiabats minimize the misfit to reference models. We follow the method of Bina and Silver [1990], except that we minimize the misfit to seismological  $\rho(z)$  and  $T_0^{\Phi}$  rather than to  $\rho(z)$  and  $V_{\Phi}(z)$ .

Thus, rather than fit a  $V_{\Phi}(z)$  profile resulting from inversion of seismic data for independent  $V_P(z)$  and  $V_S(z)$  profiles, we directly solve the forward problem for  $T_0^{\Phi}$  for each candidate composition and compare those solutions to  $T_0^{\Phi}$  from our observations of core reflected phases. Our approach thus avoids comparison with problematic  $V_{\Phi}(z)$  profiles, and it obviates the necessity of attempting to modify an AK135  $V_{\Phi}(z)$  profile to yield a larger  $T_0^{\Phi}$  value—modification which would result in a multitude of nonunique models, many of which would violate other seismological constraints. Furthermore, our methodology allows us to directly vary the

reference  $T_0^\Phi$  between different models so as to determine the effect upon mantle compositional constraints of assigning responsibility for our observed delays to the lower mantle or to shallower regions. We present two end-member interpretations in Figure 2: in the bottom panel (AK135+3.5s), all 3.5 s of  $T_0^\Phi$  delay is assumed to occur in the lower mantle; in the top panel (AK135), it is all mapped into the crust and upper mantle.

While fitting  $\rho(z)$  provides no support for an Fe-enriched lower mantle, our measured mantle  $T_0^\Phi$  of AK135+3.5s results in best-fitting lower mantle compositions which are less enriched in Si than those fit to

the  $T_0^\Phi$  calculated for AK135 (Figure 2). The high- $\gamma_{D0}$  equations of state for ferromagnesian silicate perovskite [Bina, 1995] which had supported a Si-enriched lower mantle are now consistent with less Si-rich compositions, nearer to pyrolite [Ringwood, 1975]. The low- $\gamma_{D0}$  equations of state continue to show no evidence for lower mantle Si-enrichment; indeed, these latter are even consistent with Si-depletion for low estimates of lower mantle temperature.

While the radial  $V_S$  model TIP for the Tibetan Plateau indicates that upper mantle  $V_S$  is slow [Zhao *et al.*, 1991], recent 3-D tomographic imaging of Tibetan structure reveals fast  $V_S$  and slow  $V_P$  in the upper mantle beneath our stations [J. VanDecar, pers. comm.]. Furthermore,  $SS - S$  residuals for the Tarim Basin reveal slow  $V_S$  in the upper mantle beneath our Lop Nur source [S. Roecker and I. Dricker, pers. comm.]. Thus, while we cannot definitively partition the observed  $T_0^\Phi$  delay between upper and lower mantle structure, we obtain two models in which either all or none of the observed delay arises within the lower mantle. Our results for these two end-members demonstrate the potential utility of  $T_0^\Phi$  in constraining mantle composition and illustrate the sensitivity of such inferences to uncertainties in both seismic data analysis and mineral equations of state.

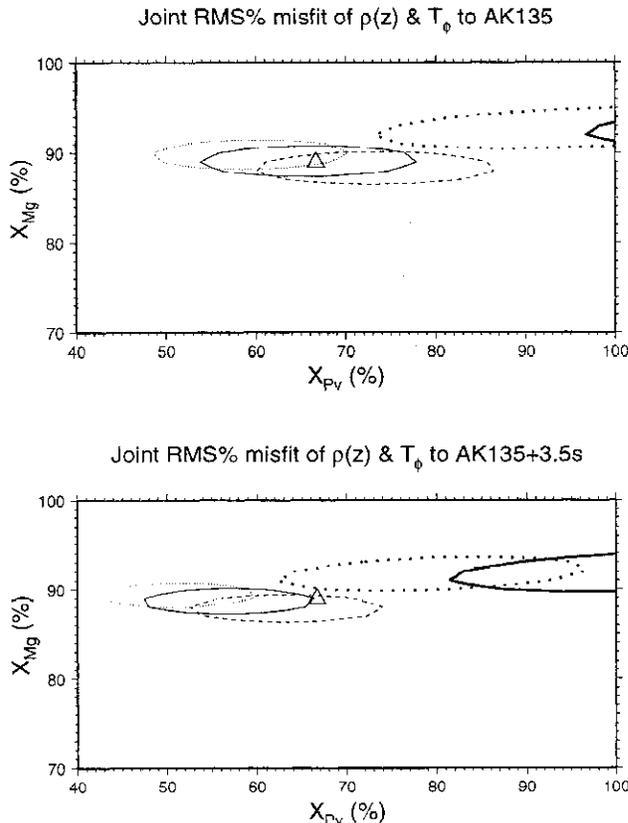
Extension of such results to other regions requires that lateral heterogeneity in  $V_\Phi$  be relatively small in the lower mantle. Tomographic imaging of mantle seismic structure yields values for  $d\ln V_S/d\ln V_\Phi$  "significantly greater" than unity [Dziewonski and Su, 1995] and values for  $d\ln V_S/d\ln V_P$  of about 2 [Woodhouse and Dziewonski, 1987], the latter rising from 1.7 to 2.6 with increasing depth in the lower mantle [Robertson and Woodhouse, 1996], thus supporting the hypothesis that lateral heterogeneity of  $V_\Phi$  is significantly less than that of  $V_S$  in the lower mantle.

Additional measurements of mantle  $T_0^\Phi$  from a growing body of seismic data, in conjunction with improving knowledge of the equations of state of mantle minerals, should place tighter constraints upon lower mantle composition and lateral heterogeneity. Furthermore, the possibility of measuring  $T_0^\Phi$  for phases which interact with transition zone seismic discontinuities may offer a direct probe into the distribution of  $T_0^\Phi$  delays between upper and lower mantle structure.

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**Figure 2.** Contours (0.5%) of joint RMS misfit to reference seismic models of calculated  $\rho(z)$  and  $T_0^\Phi$  as functions of candidate lower mantle composition, after the method of Bina and Silver [1990]. Compositions are parameterized in terms of the molar ratios  $X_{Mg} \equiv Mg/(Mg + Fe)$  and  $X_{Pv} \equiv Si/(Mg + Fe)$ . Triangle denotes pyrolite composition. Heavy contours are for high- $\gamma_{D0}$  perovskite equation of state (line 1, Table 1 of Bina [1995]); light contours are for low- $\gamma_{D0}$  equation of state (line 2, loc. cit.). Adiabatic temperatures at 660 km are 1800 K (dotted), 2000 K (solid), or 2200 K (dashed). In the top panel,  $\rho(z)$  and  $T_0^\Phi$  are fit to AK135 values. Thus, assuming a low- $\gamma_{D0}$  perovskite equation of state and a temperature of 2000 K at 660 km, lower mantle compositions which best fit AK135 densities and travel time fall within the solid light contour, overlapping pyrolite. The absence of a heavy dashed contour indicates that for a hot (2200 K at 660 km) lower mantle, even Si-enrichment to 100% perovskite cannot match seismic data for high- $\gamma_{D0}$  perovskite equations of state. In the bottom panel,  $\rho(z)$  is again fit to AK135, but  $T_0^\Phi$  is fit to AK135+3.5s to match our observed travel times.

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