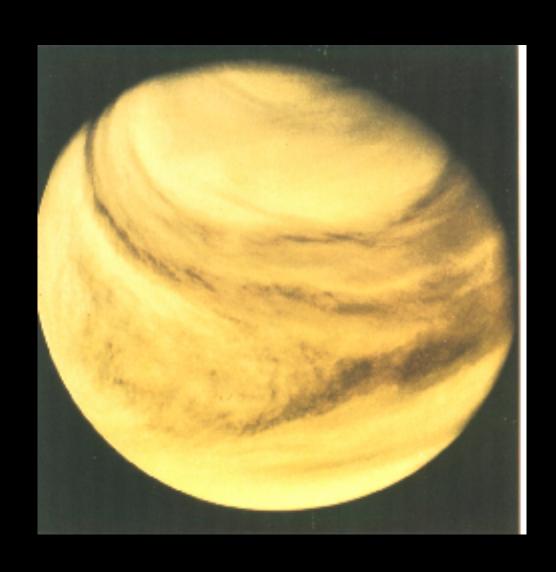
## COMPARATIVE TECTONICS OF VENUS AND EARTH

Donna M. Jurdy

Department of Earth & Planetary Sciences
Northwestern University

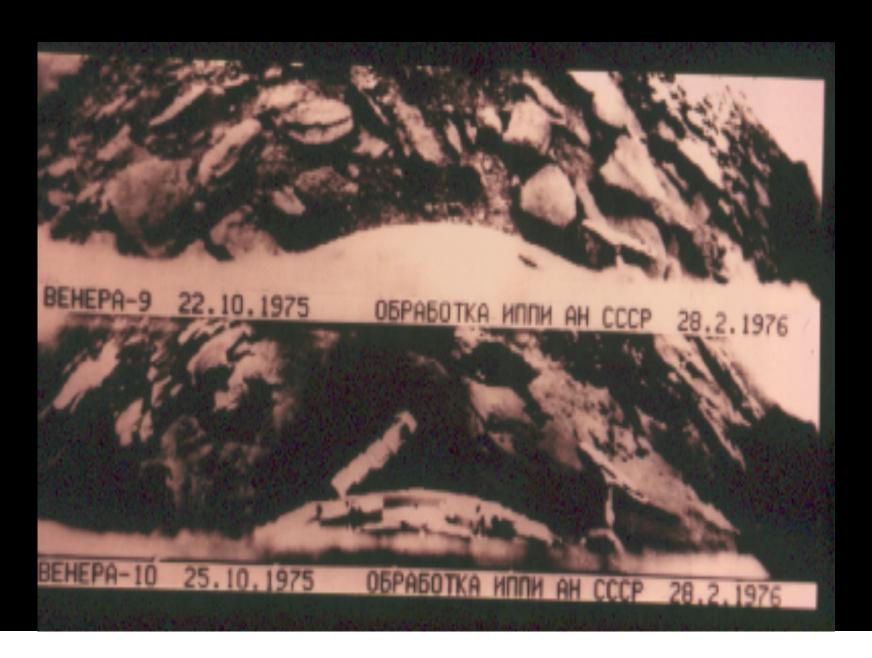
## Pioneer image of Venus



## 50's sci-fi view of Venus



## Venera 9 - 1975



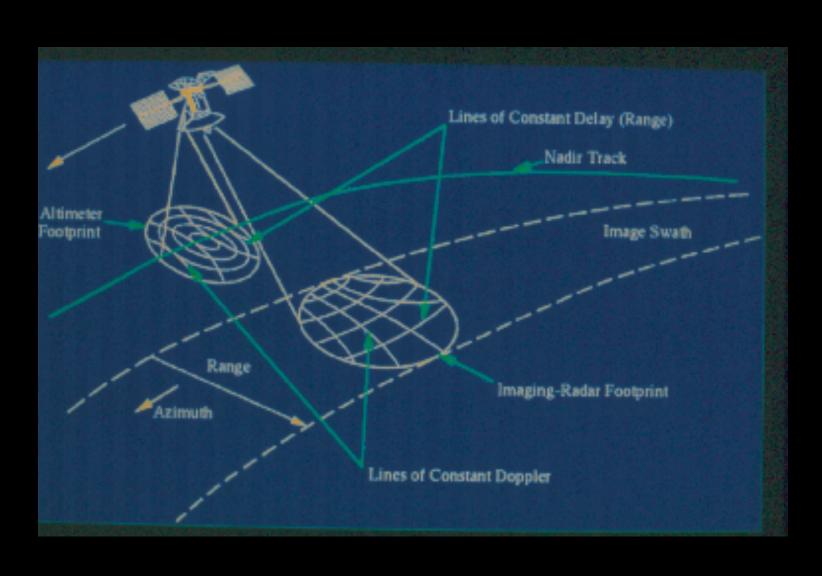
## Venus Magellan Mission



## Magellan Antenna

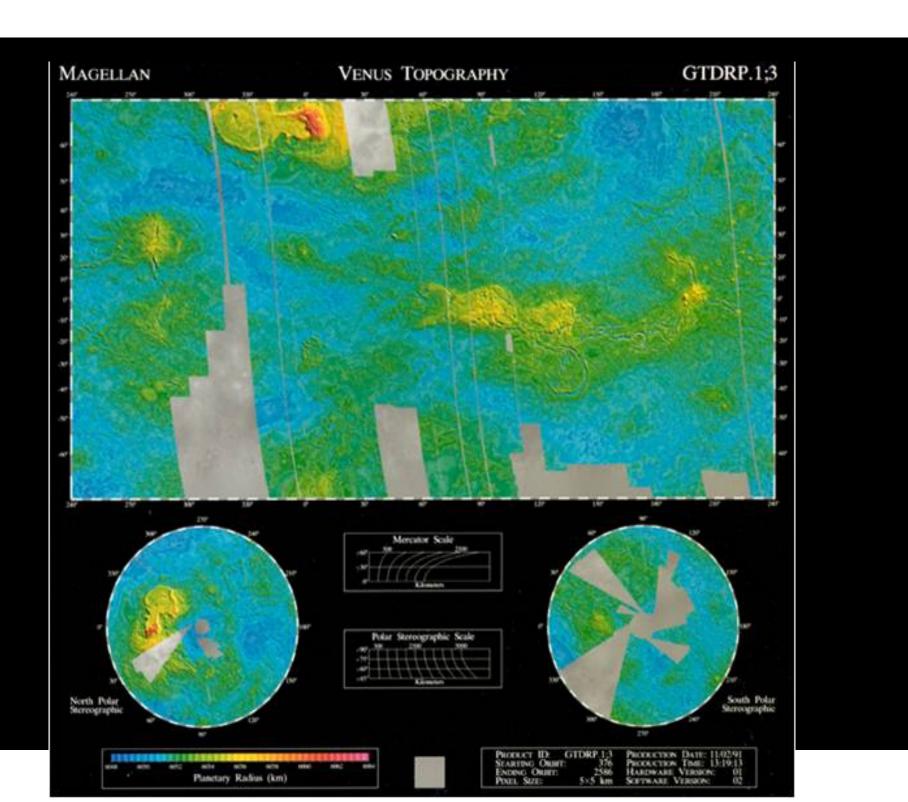


## Magellan Radar Imaging

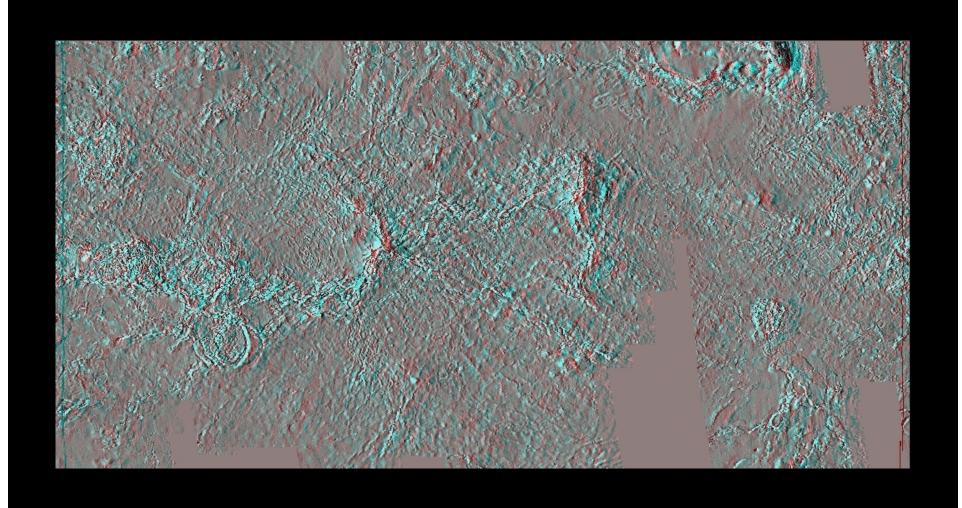


## Venus- radar mosaic

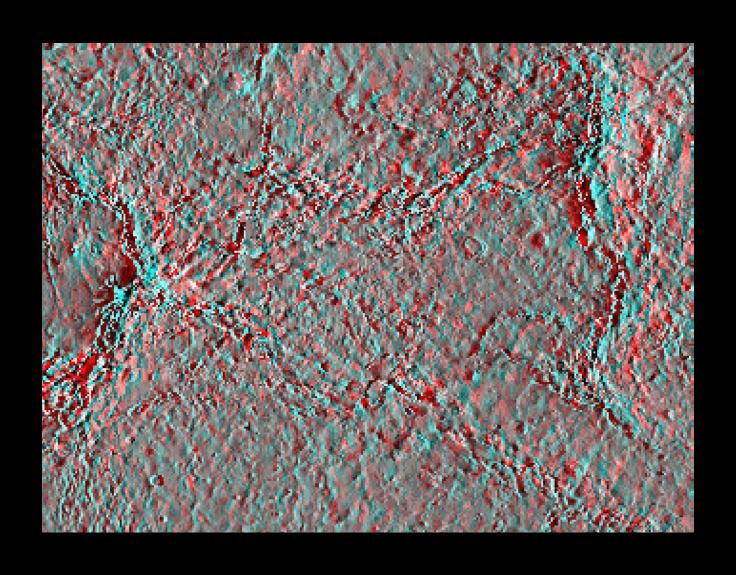




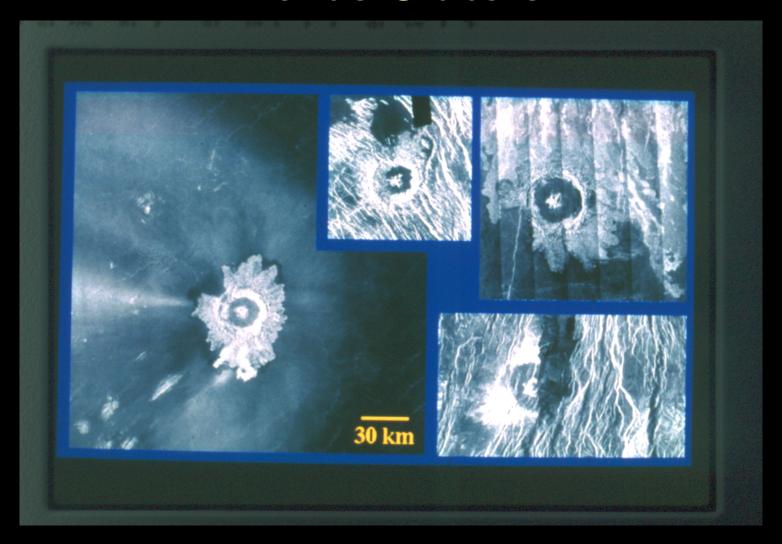
## Venus 3D Topo



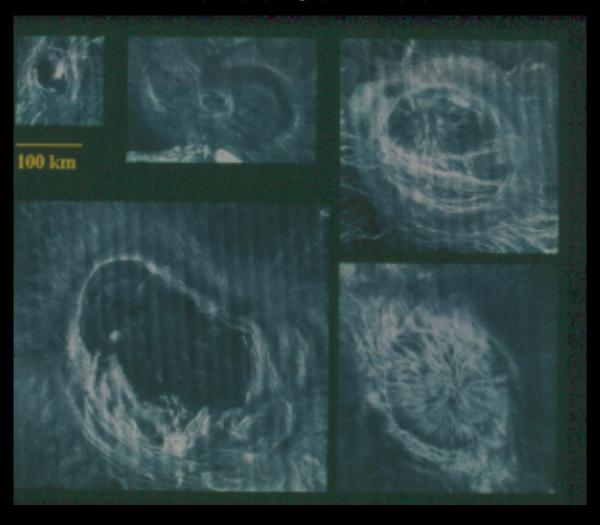
## BAT 3D Topo



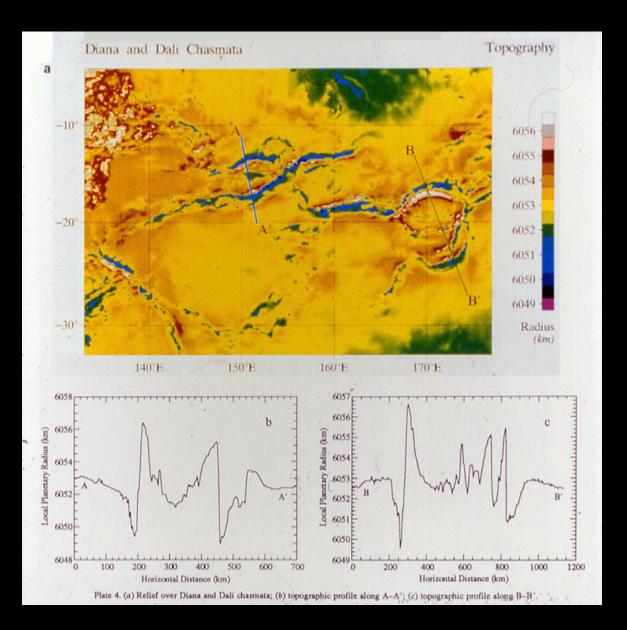
## Venus Craters



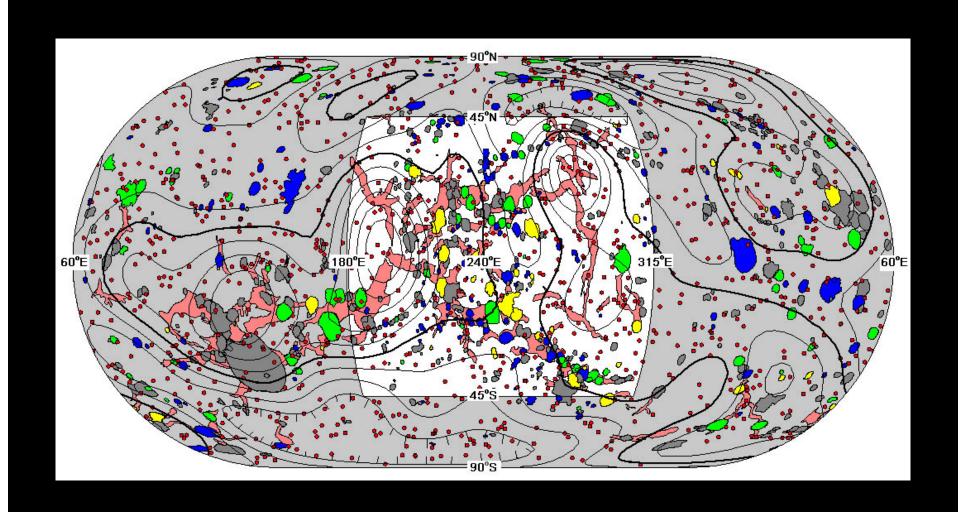
## Venus Coronae



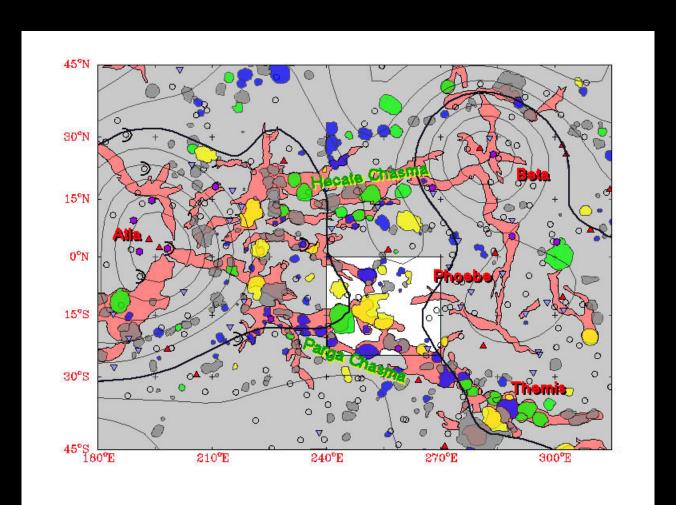
## Venus Chasmata



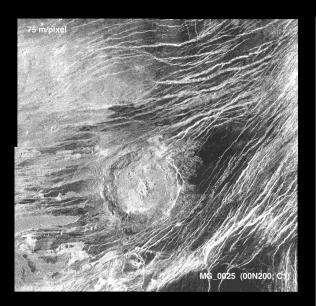
## Venus Chasmata, Coronae, Craters & Geoid

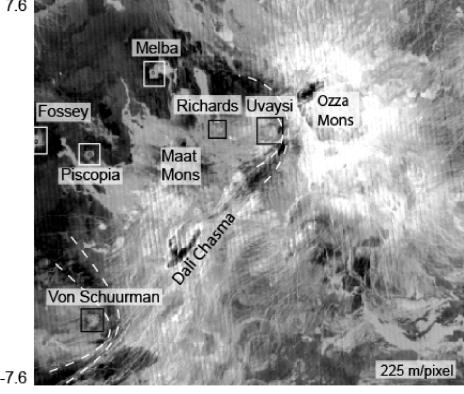


### Venus BAT Region Chasmata, Craters, Coronae & Geoid



## Crater Uvaysi Uvaysi (2.3 N, 198.2 E, 38.0 km)





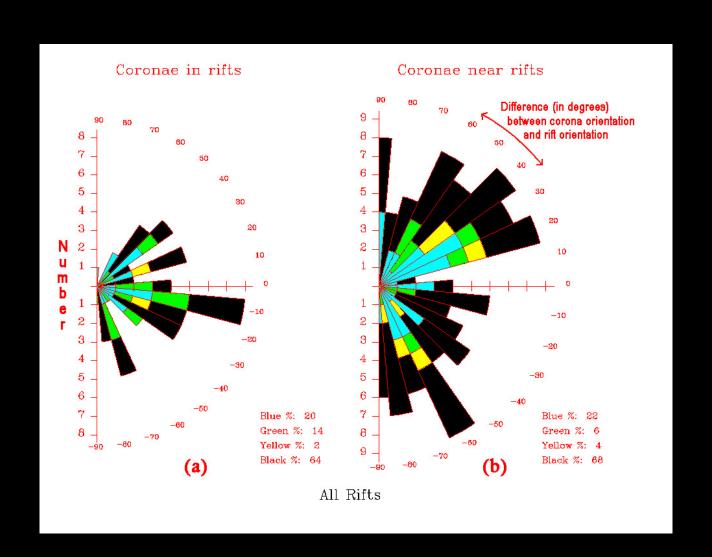


188.6 206.2

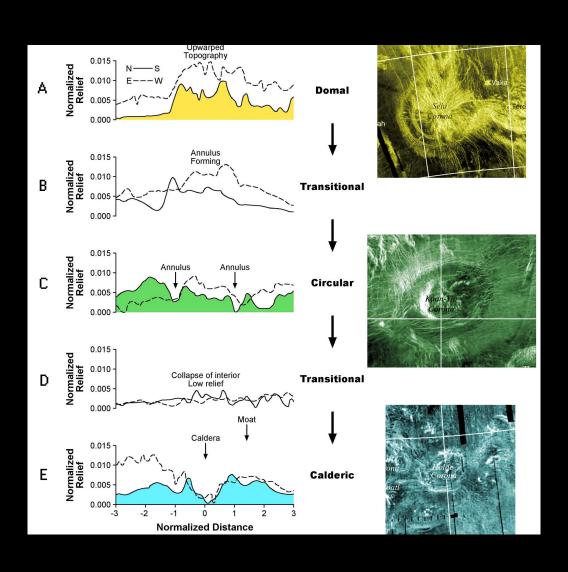
### Parga Chasm with coronae



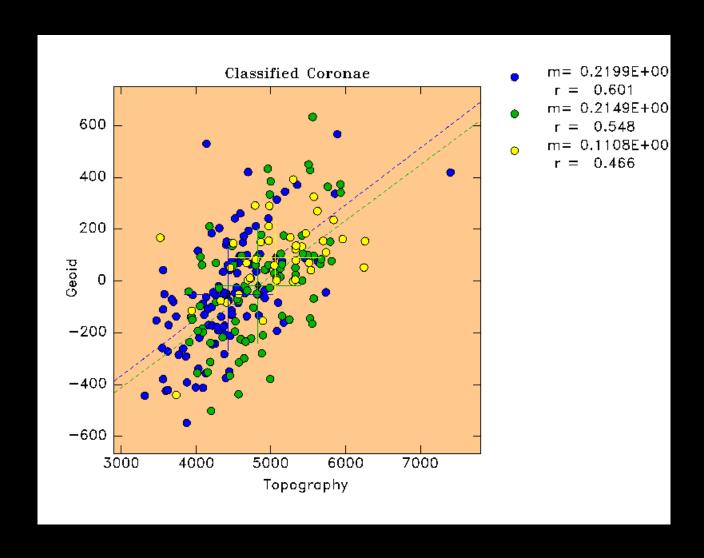
## Coronae orientation



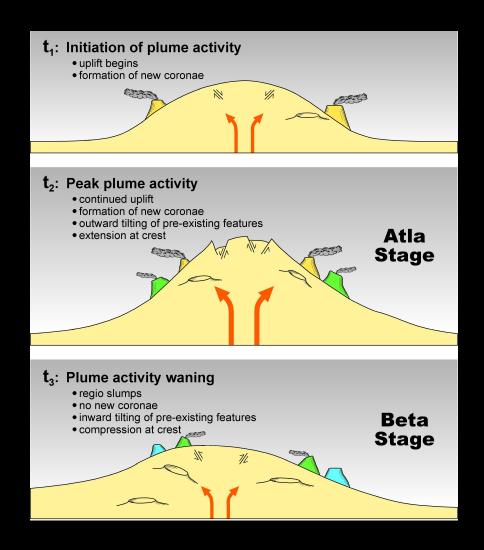
## Coronae Classification

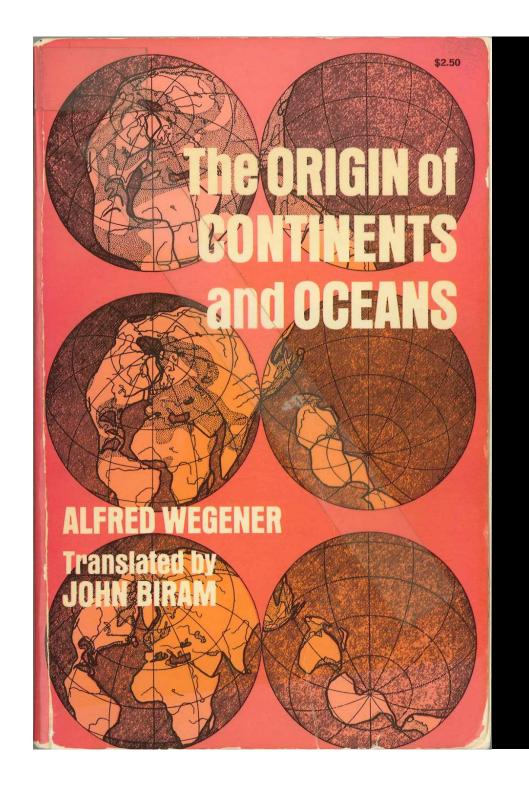


## Coronae by type



## Regio Stages

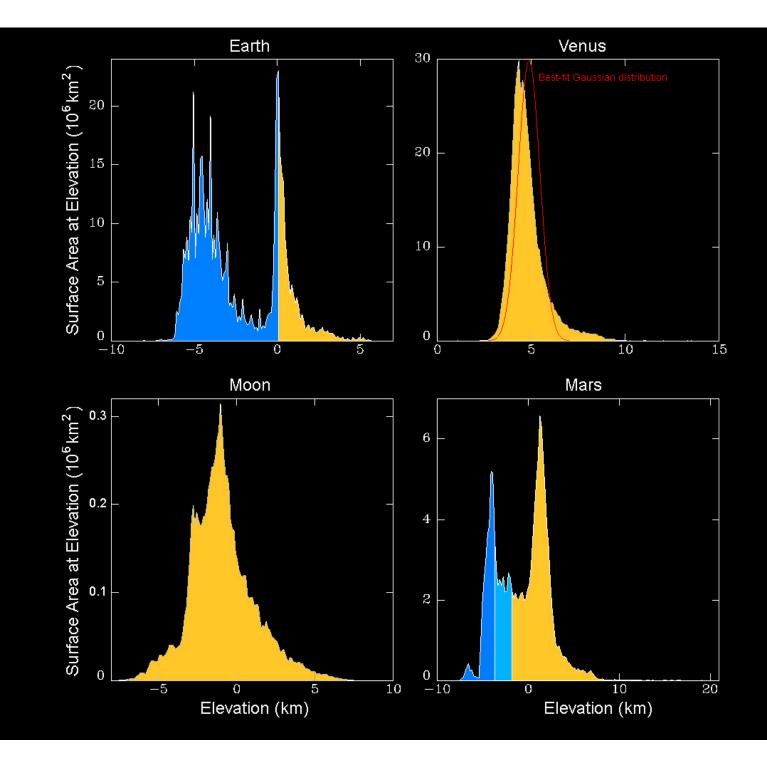




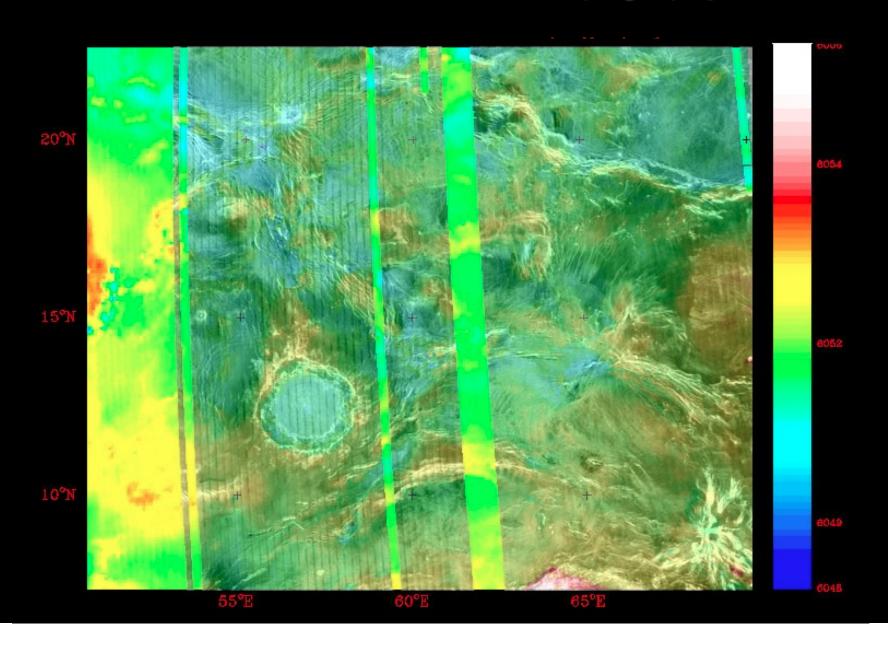
"In the whole of geophysics there is hardly another law of such clarity and reliability as this - that there are two preferential levels for the world's surface which occur in alternation side by side, and are represented by the continents and ocean floors, respectively. It is therefore very surprising that scarcely anyone has tried to explain this law, which has, after all, been well known for some time... In this way we have achieved for the first time a plausible explanation..."

## - Alfred Wegener

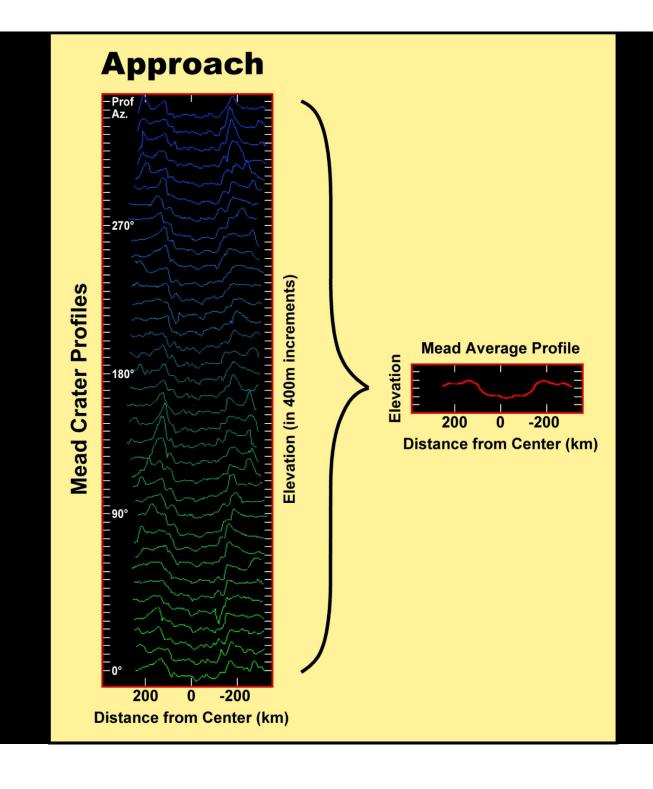
## Comparative Hypsometry



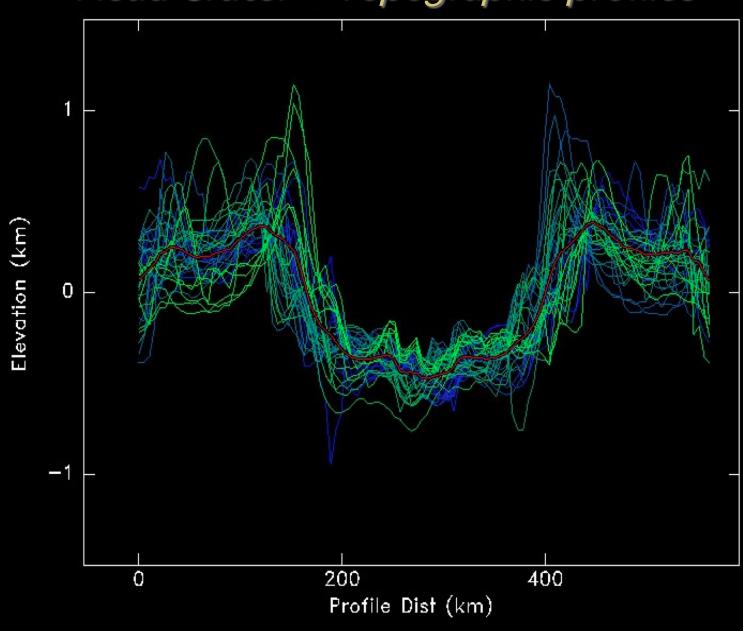
## Mead Crater - Radar & Topography



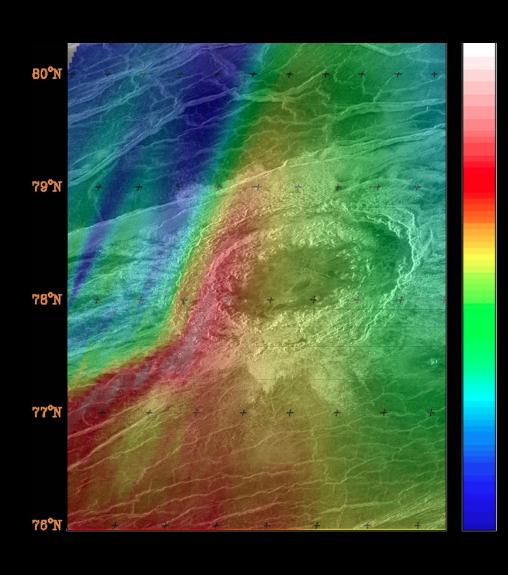
## **Average Profiling**



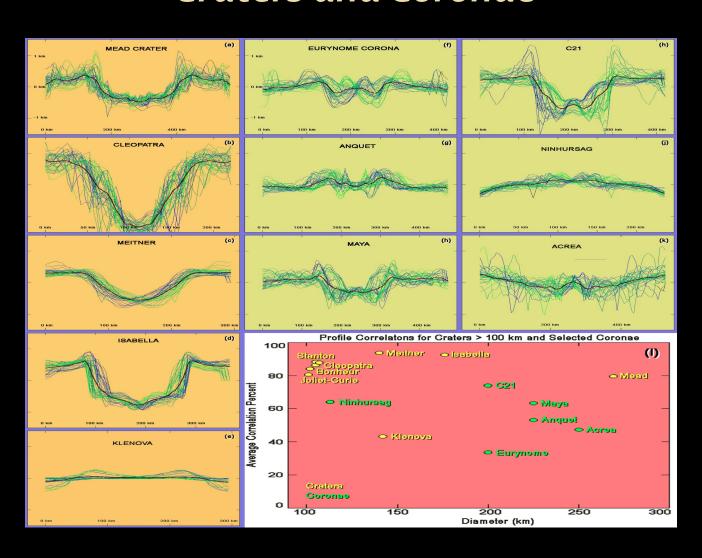
## Mead Crater - Topographic profiles



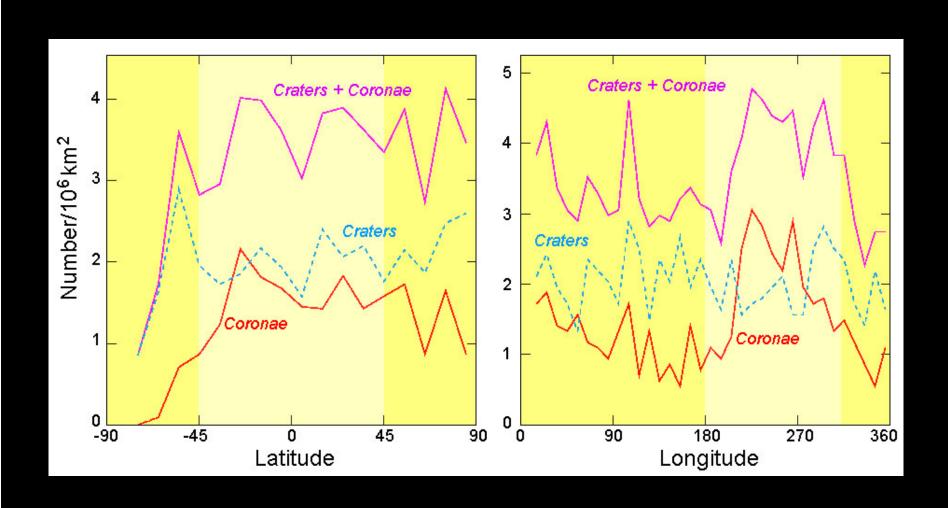
## Klenova – crater or corona?



## Topography Correlation Craters and Coronae



## Feature Densities



## Plumes- a controversial topic!

Parties and Planes

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Edited by G.R. Foulger and D.M. Jurdy

Plates, Plumes, and Planetary Processes





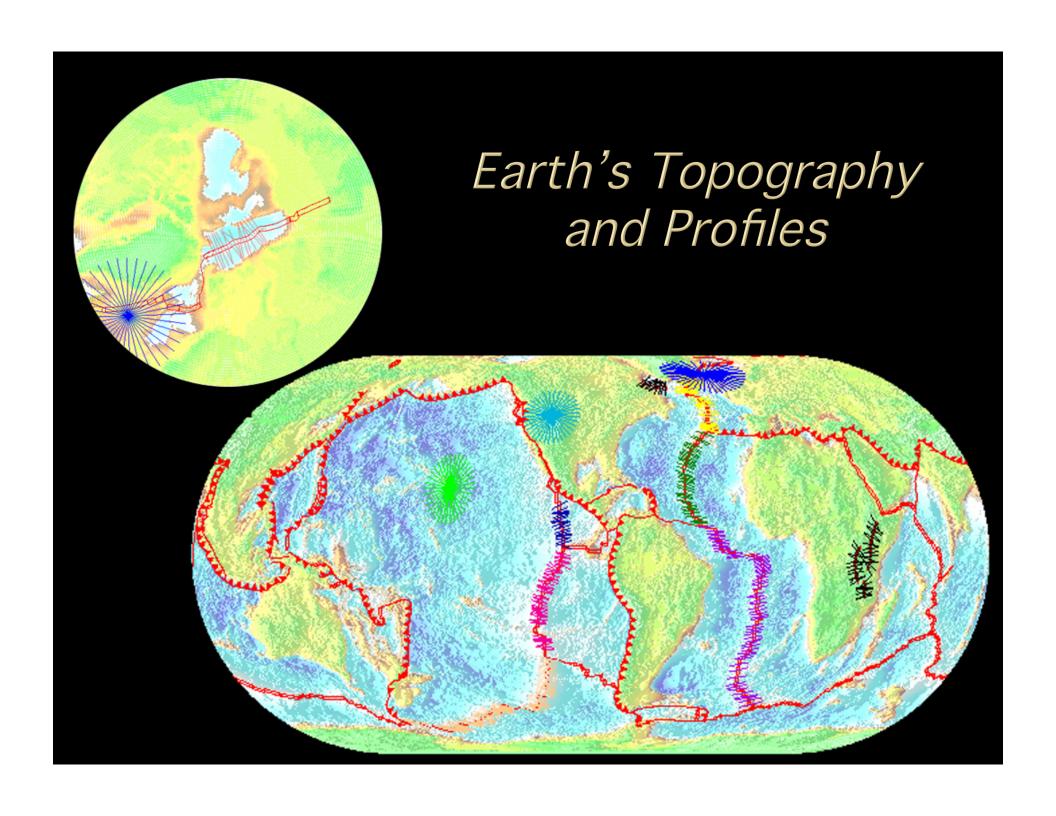
## Plates, Plumes, and Planetary Processes



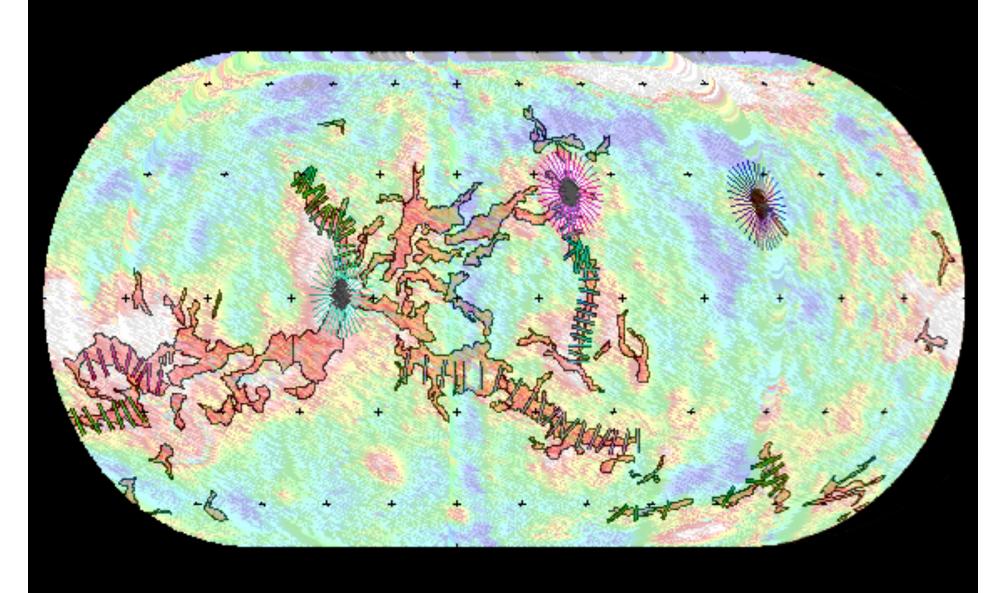
edited by Gillian R. Foulger and Donna M. Jurdy



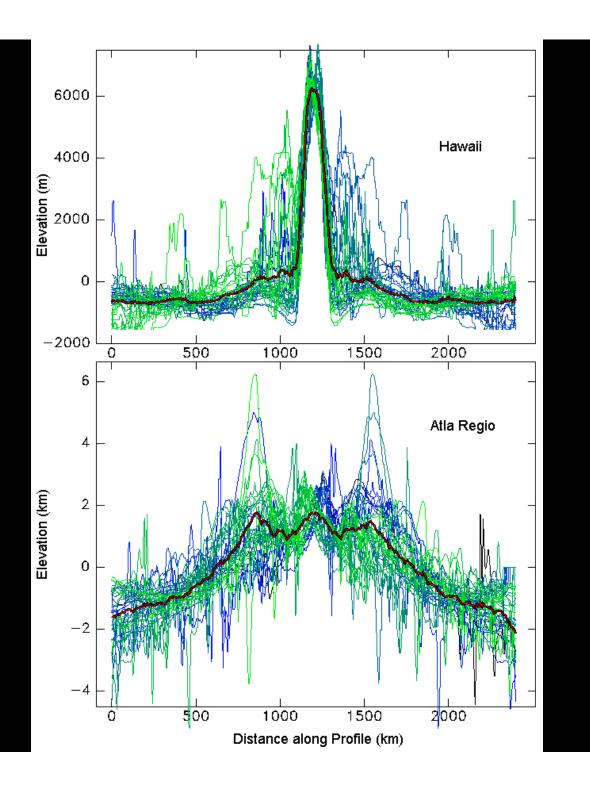
978-0-8137-2430-0

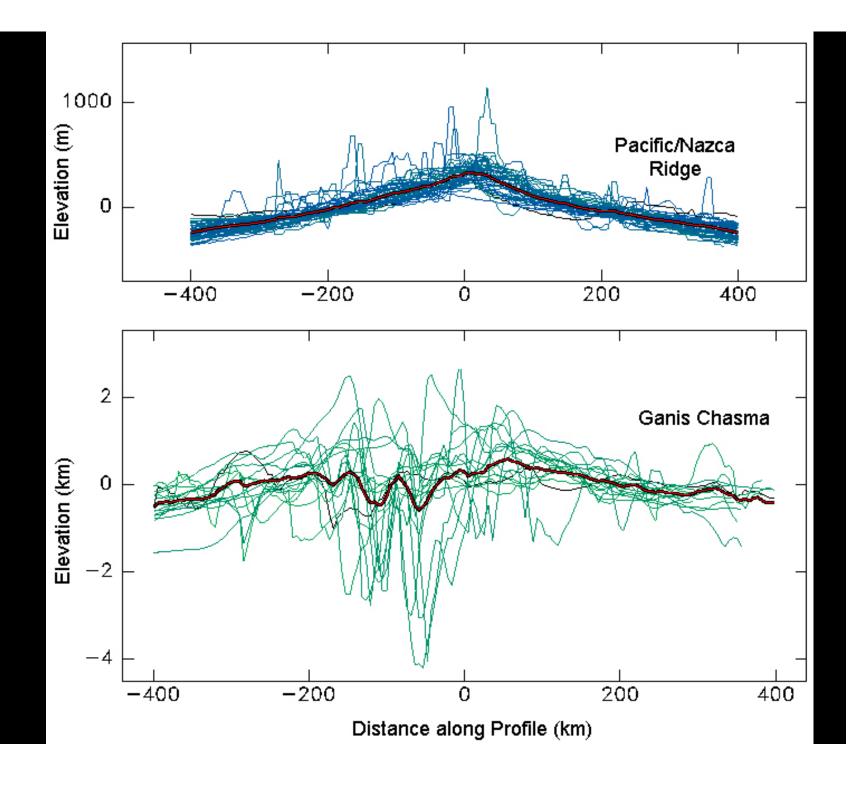


## Venus Topography and Profiles

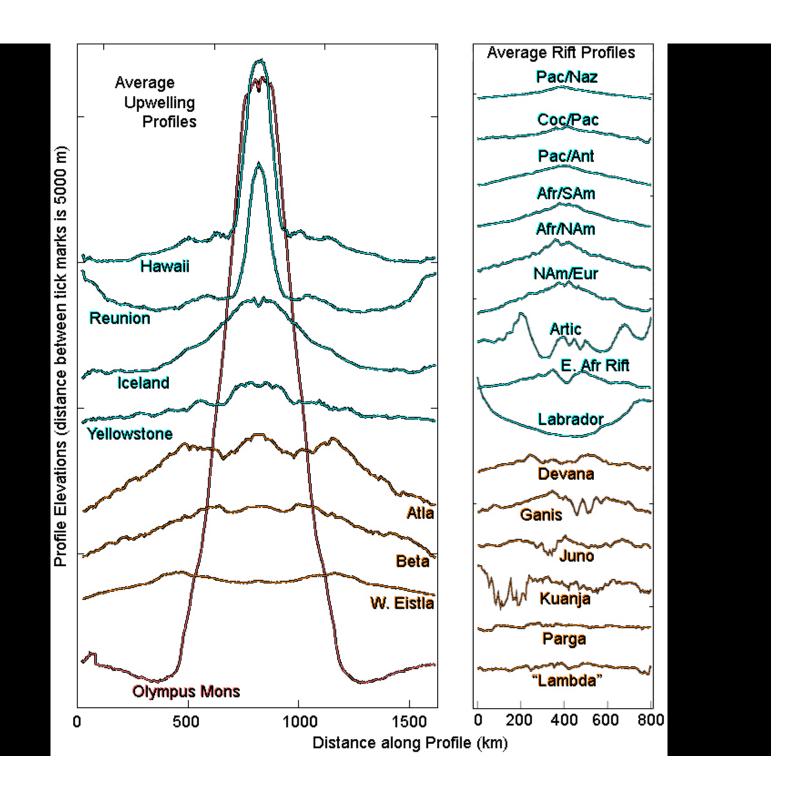


# Earth/Venus Comparison: Uplifts

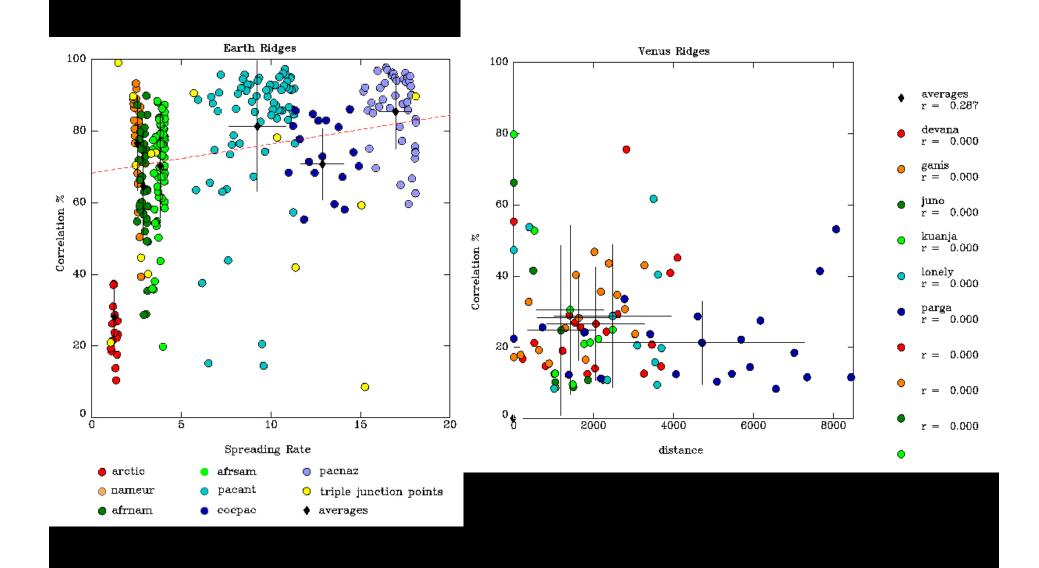




## Average Profiles: Earth and Venus



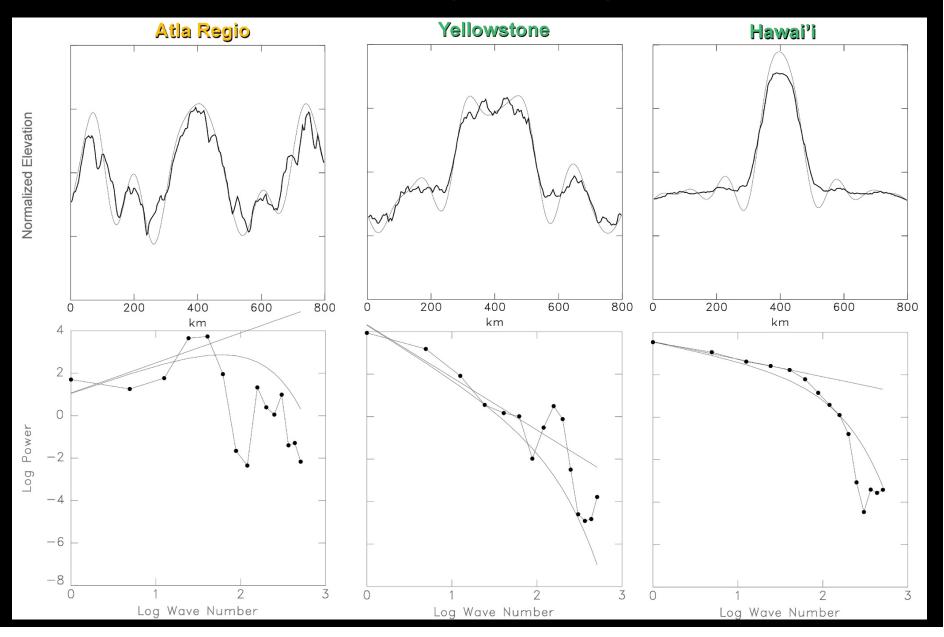
## Correlation Comparisons



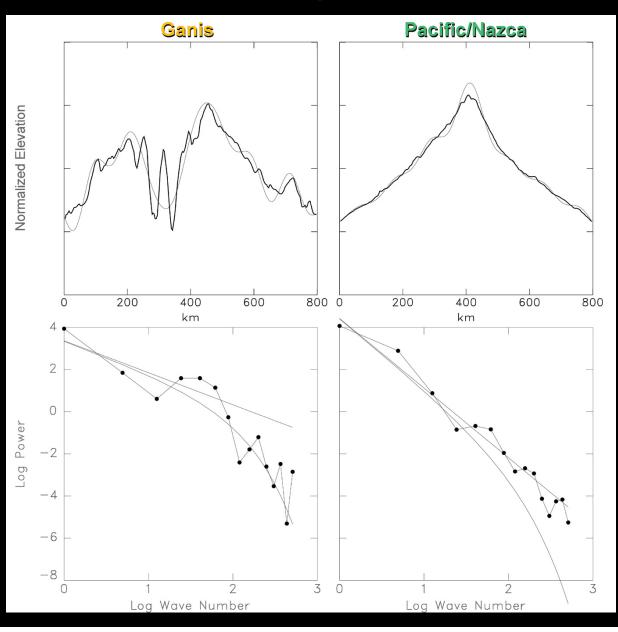
## Uplift Correlations

	hawaii	reunio	icelan	yellow	atla	beta	w.eist	olymon	
Hawaii	1.000	0.592	0.765	0.727	0.567	0.525	0.352	0.851	hawaii
Reunion	0.592	1.000	0.412	-0.034	-0.278	-0.339	-0.507	0.403	reunion
Iceland	0.765	0.412	1.000	0.693	0.578	0.530	0.295	0.936	iceland
Yellowstone	0.727	-0.034	0.693	1.000	0.884	0.900	0.801	0.809	y'stone
Atla	0.567	-0.278	0.578	0.891	1.000	0.968	0.938	0.600	atla
Beta	0.525	-0.340	0.530	0.907	0.968	1.000	0.952	0.609	beta
W. Eistla	0.352	-0.508	0.295	0.808	0.938	0.952	1.000	0.372	w.eistla
Oly. Mons	0.851	0.403	0.936	0.809	0.600	0.608	0.372	1.000	olymons
	hawaii	reunio	icelan	yellow	atla	beta	w.eist	olymon	

## Fourier Analyses (Uplifts)



## Fourier Analyses (Rifts)



## Conclusions

- Cross -correlation of like features from Earth yields closest matches
- Between Earth and Venus, rift features provide closest match. Of terrestrial hotspots, Yellowstone most closely matches Venus' regiones.
- Topography of local constructs of the regiones is dominated by rifting, but the longer-wavelength profiles reflect the larger-area upwelling processes.
- If Venus' chasmata are analogous to terrestrial spreading centers, spreading on Venus is very slow.

## Inner Solar System

Radius (km)	2439	6052	6378	1738	3398
Mass (kg)	3.30x10 <sup>23</sup>	4.87x10 <sup>24</sup>	5.98x10 <sup>24</sup>	7.35x10 <sup>22</sup>	6.42x10 <sup>23</sup>
Density (kg/m <sup>3</sup> )	5420	5250	5520	3340	3940
Distance from the Sun (A.U)	0.387	0.723	1.000		1.524

## Venus- radar mosaic

